

Thermal radiation effect on an unsteady hydromagnetic free convective oscillatory Couette flow of a viscous fluid embedded in a porous medium

S Harisingh Naik¹, K. Rama Rao² & M V Ramana Murthy³

^{1,3}Department of Mathematics & Computer Science, University College for Sciences, Osmania university, Hyderabad-7.

²Department of Mathematics, Chaithanya Bharathi Institute of Technology (C. B. I. T), Gandipet, Hyderabad, 500075, Andhra Pradesh, India.

Abstract: *This paper investigates the effect of thermal radiation on an unsteady magnetohydrodynamic free convective oscillatory Couette flow of an optically, viscous thin fluid bounded by two horizontal porous parallel walls under the influence of an external imposed transverse magnetic field embedded in a porous medium. The fluid is considered to be a gray, absorbing – emitting but non – scattering medium and the Rosseland approximation is used to describe the radiative heat flux in the energy equation. The non – dimensional governing coupled equations involved in the present analysis are solved by an efficient, accurate, and extensively validated and unconditionally stable finite difference scheme of the Crank Nicolson method and the expressions for velocity, temperature, Skin friction and rate of heat transfer has been obtained. Numerical results for velocity and temperature are presented graphically and the numerical values of Skin friction and Nusselt number have been tabulated. The effect of different parameters like thermal Grashof number, Magnetic field (Hartmann number), Prandtl number, Porosity parameter and Thermal radiation parameter on the velocity, temperature, Skin friction and Nusselt number are discussed.*

Keywords: Thermal radiation, MHD, Free convection, Couette flow, Periodic wall temperature and Finite difference method.

1. Introduction

Magnetohydrodynamic (MHD) free convection of a viscous incompressible fluid along a vertical wall in porous medium must be studied if we are to understand the behaviour of fluid motion in many applications as for example, in MHD electrical power generation, geophysics, astrophysics, etc. The problem of free convection flows of viscous incompressible fluids past a semi – infinite vertical wall has received a great deal of attention in recent years because of its many practical applications, such as in electronic components, chemical processing equipment, etc. Magneto hydrodynamic free convection flow of an electrically conducting fluid in different porous geometries is of considerable interest to the technical field due to its frequent occurrence in industrial, technological and geothermal applications. As an example, the geothermal region gases are electrically conducting and undergo the influence of magnetic field. Also, it has applications in nuclear engineering in connection with reactors cooling. The interest in this field is due to the wide range of

applications in engineering and geophysics, such as the optimization of the solidification processes of metals and metal alloys, the study of geothermal sources, the treatment of nuclear fuel debris, the control of underground spreading of chemical wastes and pollutants and the design of MHD power generators. Many papers concerned with the problem of MHD free convection flow in porous media have been published in the literature.

In view of its wide applications, the unsteady free convective MHD flow with heat transfer past a semi – infinite vertical porous moving plate with variable suction has been studied by Kim [1]. Singh and Thakur [2] have given an exact solution of a plane unsteady MHD flow of a non – Newtonian fluid. Sharma and Preeek [3] explained the behaviour of steady free convective MHD flow past a vertical porous moving surface. Singh and his co – workers [4] have analysed the effect of heat and mass transfer in MHD flow of a viscous fluid past a vertical plate under oscillatory suction velocity. Makinde et al. [5] discussed the unsteady free convective flow with suction on an accelerating porous plate. Sarangi and Jose [6] studied the unsteady free convective MHD flow and mass transfer past a vertical porous plate with variable temperature. Das and his associates [7] estimated the mass transfer effects on unsteady flow past an accelerated vertical porous plate with suction employing finite difference analysis. Das et al. [8] investigated numerically the unsteady free convective MHD flow past an accelerated vertical plate with suction and heat flux. Das and Mitra [9] discussed the unsteady mixed convective MHD flow and mass transfer past an accelerated infinite vertical plate with suction. Recently, Das and his co – workers [10] analyzed the effect of mass transfer on MHD flow and heat transfer past a vertical porous plate through a porous medium under oscillatory suction and heat source. Gireesh Kumar et al. [11] investigated effects of chemical reaction and mass transfer on MHD unsteady free convection flow past an infinite vertical plate with constant suction and heat sink. The effect of radiation is quite significant at high temperature. Radiative convective flows are encountered in countless industrial and environment processes, particularly in astrophysical studies and space technology.

Radiative heat and mass transfer play an important role in manufacturing industries for the design of fins, steel rolling, nuclear power plants, gas turbines and various propulsion devices for aircraft, combustion and furnace design, materials processing, energy utilization, temperature measurements, remote sensing for astronomy and space exploration, food processing and cryogenic engineering, as well as numerous agricultural, health and military applications. Arpaci [12] studied the interaction between thermal radiation and laminar convection of heated vertical plate in a stagnant radiating gas. England and Emery [13] have studied the thermal radiation effects of an optically thin gray gas bounded by a stationary vertical plate. In all above studies, the stationary plate was considered. Singh [14] studied the effects of Coriolis as well as magnetic force on the flow field of an electrically conducting fluid past an impulsively started infinite vertical plate. Bestman and Adjepong [15] studied the magnetohydrodynamic free convection flow, with radiative heat transfer, past an infinite moving plate in rotating incompressible, viscous and optically transparent medium. Das et al. [16] have analysed radiation effects on flow past an impulsively started infinite isothermal vertical plate. Raptis and Perdakis [17] considered the effects of thermal radiation and free convection flow past a moving vertical plate. The governing equations were solved analytically.

The convection problem in a porous medium has important applications in geothermal reservoirs and geothermal extractions. The process of heat and mass transfer is encountered in aeronautics, fluid fuel nuclear reactor, chemical process industries and many engineering applications in which the fluid is the working medium. The wide range of technological and industrial applications has stimulated considerable amount of interest in the study of heat and mass transfer in convection flows. The radiation effect on heat transfer over a stretching surface has been studied by Elbashbeshy [18]. Takharet al. [19] studied the radiation effects on MHD free convection flow of a gas past a semi-infinite vertical plate. Thermal radiation and buoyancy effects on MHD free convective heat generating flow over an accelerating permeable surface with temperature dependent viscosity has been studied by Seddeek [20]. Ghaly and Elbarbary [21] have investigated the radiation effect on MHD free convection flow of a gas at a stretching surface with a uniform free stream. In all the above studies, only steady state flows over semi-infinite vertical plate have been considered. The unsteady free convection flows over a vertical plate has been studied by Gokhale [22] and Muthucumaraswamy and Ganesan [23]. Cheng and Teckchandani [24] obtained numerical solutions for the convective flow in a porous medium bounded by two isothermal parallel plates in the presence of the withdrawal of the fluid. In all the above mentioned studies, the effect of porosity, permeability and the thermal resistance of the medium is ignored or treated as constant. However, porosity measurements by Benenati and Broseow [25] show that porosity is not constant but varies from the

surface of the plate to its interior to which as a result permeability also varies. In case of unsteady free convective flows, the combined effect of buoyancy forces from thermal and mass diffusion on forced convection was studied by Chen et al. [26]. Bejan and Khair [27] have investigated the vertical free convective boundary layer flow embedded in a porous medium resulting from the combined heat and mass transfer. Lin and Wu [28] were analysed the problem of simultaneous heat and mass transfer with the entire range of buoyancy ratio for most practical and chemical species in dilute and aqueous solutions. Rushi Kumar and Nagarajan [29] studied the mass transfer effects of MHD free convection flow of an incompressible viscous dissipative fluid past an infinite vertical plate. Mass transfer effects on

Free convection flow of an incompressible viscous dissipative fluid have been studied by Manohar and Nagarajan [30].

In fluid dynamics, Couette flow refers to the free convection flow of a viscous fluid in the space between two parallel plates, one of which moving relative to the other. The flow is driven by virtue of viscous drag force acting on the fluid and the applied pressure gradient parallel to the plates. This type of flow is named in honour of Maurice Marie Alfred Couette, a professor of physics at the French university of Angers in the late 19th century. Couette flow is frequently used in undergraduate physics and engineering courses to illustrate shear-driven fluid motion. Some important application areas of Couette motion are magnetohydrodynamics power generators and pumps, polymer technology, petroleum industry and purification of crude oil etc. An analysis of flow of Couette flow has been studied extensively for the case of horizontal channel. Choi et al. [31] studied the buoyancy effects in plane Couette flow heated uniformly from below with constant heat flux. Attia and Sayed – Ahmed [32] investigated the problem of the effect Hall currents on unsteady MHD Couette flow and heat transfer of a Bingham fluid with suction and injection. The effectiveness of variation in the physical variables on the generalized Couette flow with heat transfer in a porous medium studied by Attia [33]. Makinde and Osalusi [34] considered the problem of MHD steady flow in a channel filled with porous material with slip at the boundaries, while, Narahari [35] studied the effects of thermal radiation and free convection currents on the unsteady Couette flow between two vertical parallel plates with constant heat flux at one boundary. Israel – Cooke et al. [36] discussed oscillatory magnetohydrodynamic Couette flow of a radiating viscous fluid in a porous medium with periodic wall temperature.

The object of the present paper is to analyse the effect of thermal radiation on an unsteady magnetohydrodynamic free convective oscillatory Couette flow of an optically, viscous thin fluid bounded by two horizontal porous parallel walls under the influence of an external imposed transverse magnetic field embedded in a porous medium. The fluid is considered to be a grey, absorbing – emitting but non-scattering medium and the Rosseland

approximation is used to describe the radiative heat flux in the energy equation. The non-dimensional governing coupled equations involved in the present analysis are solved by an efficient, accurate, and extensively validated and unconditionally stable finite difference scheme of the Crank Nicolson method which is more economical from computational view point. The effects of various governing parameters on the velocity, temperature, skin friction coefficient and Nusselt number are shown in figures and tables and discussed in detail. From computational point of view it is identified and proved beyond all doubts that the Crank Nicolson methods more economical in arriving at the solution and the results obtained

are good agreement with the results of Israel – Cooke et al. [36] in some special cases. In section 2, the mathematical formulation of the problem and dimensionless forms of the governing equations are established. Solution method to these equations for the flow variables are briefly examined in section 3. The results of the previous sections are discussed in section 4. In section 5, general concluding remarks of the results of the previous sections are given.

2. Mathematical formulation

We consider the unsteady Couette flow of an electrically conducting and optically thin viscous incompressible fluid in a porous medium bounded by two infinite non-conducting horizontal parallel walls under the influence of an externally applied uniform magnetic field and radiative heat flux.

We made the following assumptions:

1. The x axis is taken along the plate in the vertical upward direction and the y axis is taken normal to the plate.
2. It is assumed that the fluid has small electrical conductivity and the electromagnetic force produced is very small, so we neglect the induced magnetic field.
3. The lower wall is suddenly moved from rest with a free stream velocity, that oscillates in time about a constant mean velocity.
4. It is assumed that the temperature of the moving lower plate oscillates in time about a non-zero constant mean.
5. The induced magnetic field, Hall Effect and viscous dissipation are assumed to be negligible as the magnetic Reynolds number of the flow is taken to be very small.
6. It is assumed that there is no applied voltage which implies the absence of an electric field.
7. Assuming Boussinesq's approximation for an incompressible fluid model.

The governing equations of the flow for an optically thin medium are

Momentum Equation:

$$\rho \frac{\partial u'}{\partial t'} = \rho \frac{dU'}{dt'} + \nu \frac{\partial^2 u'}{\partial y'^2} + \rho g \beta (T' - T_h') - \left(\frac{1}{K'} + \sigma B_o^2 \right) (u' - U') \quad (1)$$

Energy Equation:

$$\rho C_p \frac{\partial T'}{\partial t'} = \kappa \frac{\partial^2 T'}{\partial y'^2} + \nu \left(\frac{\partial u'}{\partial y'} \right)^2 - \frac{\partial q_r}{\partial y} \quad (2)$$

Where u' is the flow velocity in x' – direction, U' is the free stream velocity, U_o is the mean constant free stream velocity, ν is the viscosity, ρ is the fluid density, g is the acceleration due to gravity, β is the coefficient of thermal expansion, σ is the electric conductivity of the fluid, K' is the permeability of the porous medium, κ is the thermal conductivity, C_p is the specific heat capacity at constant pressure, T' is the fluid temperature, T_w' is the temperature of the upper wall, T_h' is the temperature of the lower wall and q_r is the radiative heat flux. Also, ω' is the frequency of oscillation, $\varepsilon \ll 1$ is a small parameter and $U'(t') = 1 + \varepsilon e^{i\omega t'}$ is the freestream velocity.

The corresponding boundary conditions are

$$\left. \begin{aligned} u' &= U_o (1 + \varepsilon e^{i\omega t'}) , T' = T_w' + \varepsilon (T_w' - T_h') e^{i\omega t'} \text{ on } y' = 0 \\ u' &= 0, T' = T_h' \text{ on } y' = h \end{aligned} \right\} \quad (3)$$

The radiative heat flux term is simplified by making use of the Rosseland approximation [37] as

$$q_r = - \frac{4\bar{\sigma}}{3k^*} \frac{\partial T'^4}{\partial y'} \quad (4)$$

Here $\bar{\sigma}$ is Stefan – Boltzmann constant and k^* is the mean absorption coefficient. It is assumed that the temperature differences within the flow are sufficiently small so that T'^4 can be expressed as a linear function of T' after using Taylor's series to expand T'^4 about the free stream temperature T_h' and neglecting higher – order terms. This results in the following approximation:

$$T'^4 \cong 4T_h'^3 T' - 3T_h'^4 \quad (5)$$

Using equations (4) and (5) in the last term of equation (2), we obtain:

$$q_r = - \frac{16\bar{\sigma}T_h'^3}{3k^*} \frac{\partial T'}{\partial y'} \quad (6)$$

Introducing (6) in the equation (2), the energy equation becomes:

$$\rho C_p \frac{\partial T'}{\partial t'} = \kappa \frac{\partial^2 T'}{\partial y'^2} + \frac{16\bar{\sigma}T_h'^3}{3k^*} \frac{\partial T'}{\partial y'} + \nu \left(\frac{\partial u'}{\partial y'} \right)^2 \quad (7)$$

Introducing the following non-dimensional quantities into the equations (1) and (7):

$$\left. \begin{aligned} y &= \frac{y'}{h}, u = \frac{u'}{U_o}, U = \frac{U'}{U_o}, t = \omega t', \omega = \frac{\omega h^2}{\nu}, \theta = \frac{T' - T'_h}{T'_w - T'_h}, Gr = \frac{g \beta h^3 (T'_w - T'_h)}{\nu U_o} \\ \zeta^2 &= \frac{h^2}{K}, M^2 = \frac{\alpha_o^2 h^2}{\nu \rho}, R = \frac{\kappa \kappa^*}{4 \alpha_h^3}, Pr = \frac{\rho C_p}{\kappa}, Ec = \frac{U_o^2}{C_p (T'_w - T'_h)}, U(t) = 1 + \alpha^t \end{aligned} \right\} (8)$$

Then the equations (1) and (7), using non – dimensional quantities (8) reduce to the following non – dimensional form of equations:

$$\omega \frac{\partial u}{\partial t} = \omega \frac{\partial U}{\partial t} + \frac{\partial^2 u}{\partial y^2} + (Gr)\theta - (\zeta^2 + M^2)(u - U) \quad (9)$$

$$\omega (Pr) \frac{\partial \theta}{\partial t} = \left(\frac{3R + 4}{3R} \right) \frac{\partial^2 \theta}{\partial y^2} + (Pr)(Ec) \left(\frac{\partial u}{\partial y} \right)^2 \quad (10)$$

The corresponding boundary conditions are

$$\left. \begin{aligned} u &= 1 + \varepsilon e^{-it}, \theta = 1 + \varepsilon e^{-it} \text{ on } y = 0 \\ u &\rightarrow 0, \theta \rightarrow 0 \text{ on } y = 1 \end{aligned} \right\} (11)$$

For practical engineering applications and the design of chemical engineering systems, quantities of interest include the following Skin friction, Nusselt number and Sherwood number are useful to compute. This is given by which is written in dimensionless form as

$$\tau = \frac{\tau_w}{\rho u_w^2}, \tau_w = - \left[\nu \frac{\partial u}{\partial y} \right]_{y=0} = - \rho U_o^2 u'(0) = - \left[\frac{\partial u}{\partial y} \right]_{y=0} \quad (12)$$

The dimensionless local surface heat flux (i.e., Nusselt number) is obtained as

$$N_u(x') = - \left[\frac{x'}{(T'_w - T'_h)} \frac{\partial T'}{\partial y'} \right]_{y'=0} \text{ then}$$

$$Nu = \frac{N_u(x')}{R_{e_x}} = - \left[\frac{\partial \theta}{\partial y} \right]_{y=0} \quad (13)$$

These two are calculated by numerical differentiation using Newton’s forward Interpolation formula. During computation of the above quantities, the non – dimensional time is fixed at $t = 1.0$. The mathematical formulation of the problem is now completed. Equations (9)& (10) present a coupled nonlinear system of partial differential equations and are to be solved by using initial and boundary conditions (11). However, exact solutions are difficult, whenever possible. Hence, these equations are solved by the Crank Nicholson method.

2. Numerical Solution by Crank Nicholson Method:

Equations (9) & (10) represent coupled system of non – linear partial differential equations which are solved numerically under the initial and boundary conditions (11) using the finite difference approximations. A linearization technique is first applied to replace the non – linear terms at a linear stage, with the corrections incorporated in subsequent iterative steps until convergence is reached.

Then the Crank Nicolson implicit method is used at two successive time levels [38]. An iterative scheme is used to solve the linearized system of difference equations. The solution at a certain time step is chosen as an initial guess for next time step and the iterations are continued till convergence, within a prescribed accuracy. Finally, the resulting block tridiagonal system is solved using the generalized Thomas – algorithm [38]. Finite difference equations relating the variables are obtained by writing the equations at the midpoint of the computational cell and then replacing the different terms by their second order central difference approximations in the direction. The diffusion terms are replaced by the average of the central differences at two successive time – levels. The computational domain is divided into meshes of dimension Δt and Δy in time and space respectively as shown in figure 1. We define the variables $B = u_y$ and $L = \theta_y$ to reduce the second order differential equations (9) & (10) to first order differential equations. The finite difference representations for the resulting first order differential equations (9)& (10) take the following forms:

$$\omega \left(\frac{u_{i+1,j+1} - u_{i,j+1} + u_{i+1,j} - u_{i,j}}{2(\Delta t)} \right) = \omega (\alpha^i \alpha^j) + \left(\frac{B_{i+1,j+1} + B_{i,j+1}}{2(\Delta y)} \right) + Gr \left(\frac{\theta_{i+1,j+1} + \theta_{i,j+1} + \theta_{i+1,j} + \theta_{i,j}}{4} \right) - (\zeta^2 + M^2) \left(\frac{u_{i+1,j+1} + u_{i,j+1} + u_{i+1,j} + u_{i,j}}{4} - U \right)$$

$$\omega \left(\frac{\theta_{i+1,j+1} - \theta_{i,j+1} + \theta_{i+1,j} - \theta_{i,j}}{2(\Delta t)} \right) = \frac{1}{Pr} \left(\frac{3R + 4}{3R} \right) \left(\frac{L_{i+1,j+1} + L_{i,j+1}}{4} - (L_{i+1,j} + L_{i,j}) \right) + QZFO \quad (14)$$

Where represents the viscous dissipation term which are known from the solution of the momentum equations and can be evaluated at the midpoint of the computational cell. Computations have been made for 2.0, 2.0, 0.71, 2.0, 2.0, 0.001, 5.0 and 1.0. Grid – independence studies show that the computational domain and can be divided into intervals with step sizes 0.0001 and 0.005 for time and space respectively. The truncation error of the central difference schemes of the governing equations is . Stability and rate of convergence are functions of the flow and heat parameters. Smaller step sizes do not show any significant change in the results. Convergence of the scheme is assumed when all of the unknowns and for the last two approximations differ from unity by less than for all values of in at every time step. Less than 7 approximations are required to satisfy this convergence criteria for all ranges of the parameters studied here.

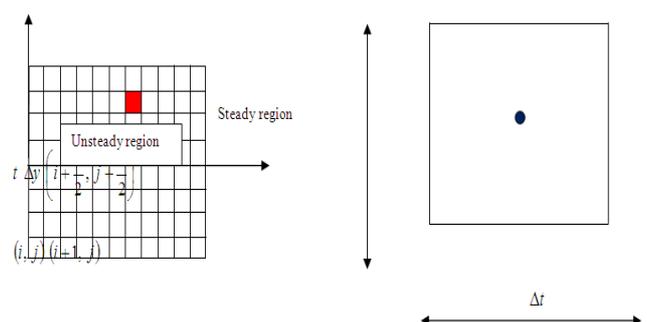


Figure 1. Mesh Layout

3. Results and Discussion:

The problem of MHD free convection and oscillatory flow of an optically thin fluid bounded by two parallel walls under the influence of an external imposed transverse magnetic field in a porous medium has been studied. By taking the radiative heat flux in the differential form and imposing an oscillatory time dependent on the coupled non-linear problem is solved for the velocity and temperature profiles. In order to understand the physical situation of the problem and hence the manifestations of the various material parameters entering the problem we have computed the numerical values of the velocity, temperature, skin friction and the Nusselt number using the software “Mat – Lab”. For the purpose of our computation, we set 2.0, 2.0, 2.0, 2.0, 0.001, 5.0 and 1.0 and Prandtl number, 0.71 which physically corresponds to the atmospheric environment (air) at fixed for the velocity and temperature profiles. The influence of the thermal Grashof number Gr on the velocity is presented in figure (2). The thermal Grashof number signifies the relative effect of the thermal buoyancy force to the viscous hydrodynamic force in the boundary layer. As expected, it is observed that there is a rise in the velocity due to the enhancement of thermal buoyancy force. Also, as Gr increases, the peak values of the velocity increases rapidly near the porous plate and then decays smoothly to the free stream velocity. Figure (3) illustrates the influences of Hartmann number M on the velocity field. It is found that the velocity decreases with increasing Hartmann number for air ($Pr = 0.71$) in presence of Hydrogen. The presence of transverse magnetic field produces a resistive force on the fluid flow. This force is called the Lorentz force, which leads to slow down the motion of electrically conducting fluid.

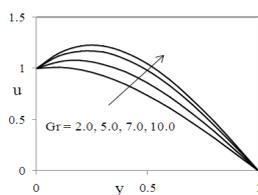


Figure 2. Velocity profiles for variations in Grashof number

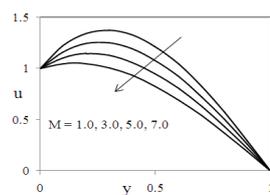


Figure 3. Velocity profiles for variations in Hartmann number

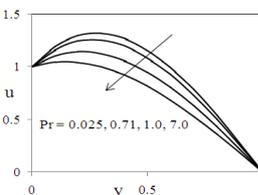


Figure 4. Velocity profiles for variations in Prandtl number

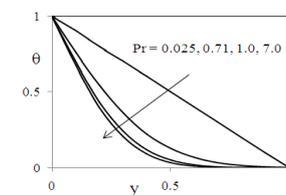


Figure 5. Temperature profiles for variations in Prandtl number

Figures (4) and (5) illustrate the velocity and temperature profiles for different values of the Prandtl number Pr . The Prandtl number defines the ratio of momentum diffusivity to thermal diffusivity. The numerical results show that the effect of increasing values of Prandtl number results in a decreasing velocity (figure (4)). From figure (5), it is observed that an increase in the Prandtl number results in a

decrease of the thermal boundary layer thickness and in general lower average temperature within the boundary layer. The reason is that smaller values of Pr are equivalent to increasing the thermal conductivities, and therefore heat is able to diffuse away from the heated plate more rapidly than for higher values of Pr . Hence in the case of smaller Prandtl numbers as the boundary layer is thicker and the rate of heat transfer is reduced.

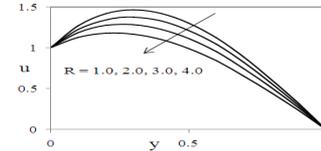


Figure 6. Velocity profiles for variations in thermal radiation parameter

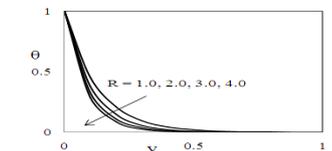


Figure 7. Temperature profiles for variations in Thermal radiation parameter

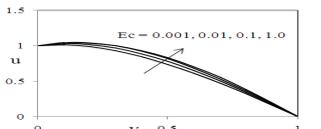


Figure 8. Velocity profiles for variations in Eckert number

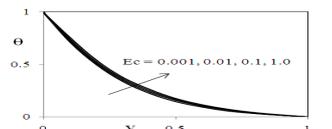


Figure 9. Temperature profiles for variations in Eckert number

The effects of the thermal radiation parameter (R) on the velocity and temperature profiles in the boundary layer are illustrated in figures (6) and (7) respectively. Increasing the thermal radiation parameter (R) produces significant increase in the thermal condition of the fluid and its thermal boundary layer. This increase in the fluid temperature induces more flow in the boundary layer causing the velocity of the fluid there to increase. The influence of the viscous dissipation parameter i.e., the Eckert number (Ec) on the velocity and temperature are shown in figures (8) and (9) respectively. The Eckert number (Ec) expresses the relationship between the kinetic energy in the flow and the enthalpy. It embodies the conversion of kinetic energy into internal energy by work done against the viscous fluid stresses. Greater viscous dissipative heat causes a rise in the temperature as well as the velocity. This behaviour is evident from figures (8) and (9). Figure (10) shows the velocity profiles for different values of the Darcy number (ζ). Clearly ζ as increases the peak value of velocity tends to decrease

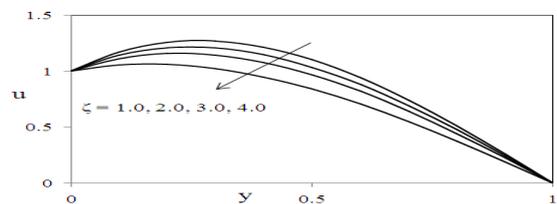


Figure 10. Velocity profiles for variations in Porosity parameter

Table – 1: Skin friction (τ) values for different values of Gr , M , Pr , ζ , R and Ec

| Gr | M | Pr | ζ | R | Ec | τ |
|------|-----|------|---------|-----|-------|--------|
| 1.0 | 1.0 | 0.71 | 1.0 | 1.0 | 0.001 | 1.8726 |
| 2.0 | 1.0 | 0.71 | 1.0 | 1.0 | 0.001 | 2.9303 |
| 1.0 | 2.0 | 0.71 | 1.0 | 1.0 | 0.001 | 0.9876 |
| 1.0 | 1.0 | 7.00 | 1.0 | 1.0 | 0.001 | 1.5504 |
| 1.0 | 1.0 | 0.71 | 2.0 | 1.0 | 0.001 | 1.3877 |
| 1.0 | 1.0 | 0.71 | 1.0 | 2.0 | 0.001 | 1.1165 |
| 1.0 | 1.0 | 0.71 | 1.0 | 1.0 | 0.100 | 1.9413 |

The profiles for Skin friction due to velocity under the effects of Grashof number, Hartmann number, Prandtl number, Porosity parameter, Thermal radiation parameter

and Eckert number are presented in the table – 1. We observe from this table – 1 the Skin friction due to velocity rises under the effects of Grashof number and Eckert number and falls under the effects of Hartmann number, Prandtl number, Porosity parameter and Thermal radiation parameter.

Table – 2: Rate of heat transfer (Nu) values for different values of Pr , Ec and R

| Pr | Ec | R | Nu |
|------|-------|-----|--------|
| 0.71 | 0.001 | 1.0 | 1.7282 |
| 7.00 | 0.001 | 1.0 | 0.9924 |
| 0.71 | 0.100 | 1.0 | 1.9287 |
| 0.71 | 0.001 | 2.0 | 1.0642 |

The profiles for Nusselt number (Nu) due to temperature profile under the effect of Prandtl number, Thermal radiation parameter and Eckert number are presented in the table– 2. We see from this table – 2, the Nusselt number due to temperature profile rises under the effect of Eckert number and falls under the effect of Prandtl number and Thermal radiation parameter. In order to ascertain the accuracy of the numerical results, the present results are compared with the existed results of Israel – Cookey *et al.* [36] for $Gr = 2.0$, $\zeta = 2.0$, $R = 2.0$, $M = 2.0$, $\omega = 5.0$, $t = 1.0$ and $Pr = 0.71$ are presented in table – 4. They are found to be in an excellent agreement.

Table – 4: Comparison of present Skin friction results (τ) with the Skin friction results (τ^*) obtained by Israel – Cookey *et al.* [36] for different values of Gr , ζ , R , M and Pr

| Gr | M | Pr | ζ | R | τ | τ^* |
|-----|-----|------|---------|-----|--------|----------|
| 1.0 | 1.0 | 0.71 | 1.0 | 1.0 | 1.7842 | 1.7840 |
| 5.0 | 1.0 | 0.71 | 1.0 | 1.0 | 2.2148 | 2.2144 |
| 1.0 | 2.0 | 0.71 | 1.0 | 1.0 | 1.4059 | 1.4051 |
| 1.0 | 1.0 | 7.00 | 1.0 | 1.0 | 1.5105 | 1.4995 |
| 1.0 | 1.0 | 0.71 | 2.0 | 1.0 | 1.5391 | 1.5377 |
| 1.0 | 1.0 | 0.71 | 1.0 | 2.0 | 1.4982 | 1.4978 |

4. Conclusions

In conclusion therefore, the effect of thermal radiation on an unsteady magnetohydrodynamic free convective oscillatory Couette flow of an optically, viscous thin fluid bounded by two horizontal porous parallel walls under the influence of an external imposed transverse magnetic field embedded in a porous medium. The non – dimensional governing coupled equations involved in the present analysis are solved by finite difference scheme of the Crank Nicolson method and the expressions for velocity, temperature, Skin friction and rate of heat transfer has been obtained. The results are presented graphically and we can conclude that the flow field and the quantities of physical interest are significantly influenced by these parameters.

1. Grashof number and Eckert number tend to retard fluid velocity whereas Hartmann number, Prandtl number, Porosity number and Thermal radiation parameter have reverse effect on it.
2. Eckert number tends to enhance fluid temperature whereas Prandtl number and Thermal radiation parameter have reverse effect on it.
3. The profiles for Skin friction due to velocity profiles rises under the effects of Grashof number and Eckert

number and falls under the effects of Prandtl number, Porosity number and Thermal radiation parameter.

4. The profiles for Nusselt number due to temperature profiles rises under the effect of Eckert number and falls under the effects of Prandtl number and Thermal radiation parameter.
5. On comparing the Skin friction results with the results of Israel – Cookey *et al.* [36] it can be seen that they agree very well.

References:

- [1]. Kim, Y. J., (2000). Unsteady MHD convective heat transfer past a semi – infinite vertical porous moving plate with variable suction, *Int. J. Engg. Sci.*, Vol. 38, pp. 833 – 845.
- [2]. Singh, B. and Thakur, C., (2002). An exact solution of plane unsteady MHD non-Newtonian fluid flows, *Ind. J. Pure Appl. Math.*, Vol. 33, No. 7, pp. 993 – 1001.
- [3]. Sharma, P. R. and Pareek, D., (2002). Steady free convection MHD flow past a vertical porous moving surface, *Ind. J. Theo. Phys.*, Vol. 50, pp. 5 – 13.
- [4]. Singh, A. K., Singh, A. K. and Singh, N. P., (2003). Heat and mass transfer in MHD flow of a viscous fluid past a vertical plate under oscillatory suction velocity, *Ind. J. Pure Appl. Math.*, Vol. 34, No. 3, pp. 429 – 442.
- [5]. Makinde, O. D., Mango, J. M. and Theuri, D. M., (2003). Unsteady free convection flow with suction on an accelerating porous plate, *AMSEJ. Mod. Meas. Cont. B.*, Vol. 72, No. 3, pp. 39 – 46.
- [6]. Sarangi, K. C. and Jose, C. B., (2005). Unsteady free convective MHD flow and mass transfer past a vertical porous plate with variable temperature, *Bull. Cal. Math. Soc.*, Vol. 97, No. 2, pp. 137 – 146.
- [7]. Das, S. S., Sahoo, S. K. and Dash, G. C., (2006). Numerical solution of mass transfer effects on unsteady flow past an accelerated vertical porous plate with suction, *Bull. Malays. Math. Sci. Soc.*, Vol. 29, No. 1, pp. 33 – 42.
- [8]. Das, S. S., Satapathy, A., Das, J. K. and Sahoo, S. K., (2007). Numerical solution of unsteady free convective MHD flow past an accelerated vertical plate with suction and heat flux, *J. Ultra Sci. Phys. Sci.*, Vol. 19, No. 1, pp. 105 – 112.
- [9]. Das, S. S. and Mitra, M., (2009). Unsteady mixed convective MHD flow and mass transfer past an accelerated infinite vertical plate with suction, *Ind. J. Sci. Tech.*, Vol. 2, No. 5, pp. 18 – 22.
- [10]. Das, S. S., Satapathy, A., Das, J. K. and Panda, J. P., (2009). Mass transfer effects on MHD flow and heat transfer past a vertical porous plate through a porous medium under oscillatory suction and heat source, *Int. J. Heat Mass Transfer*, Vol. 52, pp. 5962 – 5969.
- [11]. Gireesh Kumar, J., Satya Narayana, P. V. and Rama Krishna, S., (2009). Effects of chemical reaction and mass transfer on MHD unsteady free convection flow past an infinite vertical plate with constant suction and heat sink, *J. Ultra Scientist*, Vol. 21, No. 3, pp. 12 – 28.

- [12]. Arpaci, V.S., (1968). Effect of thermal radiation on the laminar free convection from a heated vertical plate, *Int. J. Heat Mass Transfer*, Vol. 11, pp. 871 – 881.
- [13]. England, W. G. and Emery A.F., (1969). Thermal radiation effects on the laminar free convection boundary layer of an absorbing gas, *Journal of Heat Transfer*, Vol.91, pp. 37 – 44.
- [14]. Singh, A.K., (1984). Hydromagnetic free convection flow past an impulsively started vertical infinite plate in a rotating fluid, *Int. Comm. Heat Mass Transfer*, Vol. 11, pp.399 – 406.
- [15]. Bestman, A.R. and Adjepong, S.K., (1988). Unsteady hydromagnetic free-convection flow with radiative heat transfer in a rotating fluid, *Astrophysics and Space Science*, Vol. 143, pp. 73 – 80.
- [16]. Das, U. N., Deka, R.K. and Soundalgekar, V.M., (1996). Radiation effects on flow past an impulsively started vertical infinite plate, *J.Theoretical Mechanics*, Vol. 1, pp. 111 – 115.
- [17]. Raptis, A. and Perdikis, C., (1999). Radiation and free convection flow past a moving plate, *Applied Mechanics and Engineering*, Vol. 4, pp. 817 – 821.
- [18]. Elbashbeshy, E. M. A., (2000). Radiation effect on heat transfer over a stretching surface, *Canad. J. Phys.*, Vol. 78, pp. 1107 – 1112.
- [19]. Takhar, H.S., Gorla, R.S.R. and Soundalgekar, V. M., (1996). Radiation effects on MHD free convection flow of a gas past semi – infinite vertical plate, *Int. Numer. Methods Heat Fluid Flow*, Vol. 6, pp. 77 – 83.
- [20]. Seddeek, M. A., (2001). Thermal radiation and buoyancy effects on MHD free convective heat generating flow over an accelerating permeable surface with temperature dependent viscosity, *Canad. J. Phys.*, Vol. 79, pp. 725 – 732.
- [21]. Ghaly, A.Y. and Elbarbary, E. M. E., (2002). Radiation effect on MHD free convection flow of a gas at a stretching surface with a uniform free stream, *J. Appl. Math.*, Vol. 2, pp. 93 – 103.
- [22]. Gokhale, M. Y., (1991). Magnetohydrodynamic transient free convection past a semi – infinite vertical plate with constant heat flux, *Canad. J. Phys.* 69, 1451-1453 (1991).
- [23]. Muthucumaraswamy, R. and Ganesan, P., (1998). Unsteady flow past an impulsively started vertical plate with heat and mass transfer, *Heat and Mass Transfer*, Vol.34, pp. 187 – 193.
- [24]. Cheng, P. and Teckchandani, L., (1977). Numerical solutions for transient heating and fluid withdrawal in a liquid-dominated geothermal reservoir, in the earth's crust: its nature and physical properties, Vol. 20, pp. 705 – 721.
- [25]. Benenati, R. F. and Brosilow, C. B., (1962). *A. I. Ch. E. J.*, Vol. 81, pp. 359 – 361.
- [26]. Chen T. S., Yuth, C. F. and Moutsoglou, A., (1980). Combined heat and mass transfer in mixed convection along vertical and inclined plates. *Int. J. Heat Mass Transfer*, Vol. 23, pp. 527 – 537.
- [27]. Bejan, A. and Khair, K. R., (1985). Mass transfer to natural convection boundary layer flow driven by heat transfer, *ASME J. of Heat Transfer*, Vol. 107, pp. 1979 – 1981.
- [28]. Lin, H. T. and Wu, C. M., (1995). Combined heat and mass transfer by laminar natural convection from a vertical plate, *Int. J. Heat and Mass Transfer*, Vol. 30, pp. 369 – 376.
- [29]. Rushi Kumar, B. and Nagarajan, A. S., (2007). Mass transfer effects of MHD free convection flow of an incompressible viscous dissipative fluid, *IRPAM*, Vol. 3, No. 1, pp. 145 – 157.
- [30]. Manohar, D. and Nagarajan, A. S., (2001). Mass transfer effects on free convection flow of an incompressible viscous dissipative fluid, *Journal of Energy, Heat and Mass Transfer*, Vol. 23, pp. 445 – 454.
- [31]. Choi, C. K., Chung, T. J. and Kim, M. C., (2004). Buoyancy effects in plane couette flow heated uniformly from below with constant heat flux, *International Journal of Heat and Mass Transfer*, Vol. 47, pp. 2629 – 2636.
- [32]. Attia, H. A. and Sayed – Ahmed, M. E., (2004). Hall Effect on unsteady MHD Couette flow and heat transfer of a Bingham fluid with suction and injection, *Applied Mathematical Modelling*, Vol. 28, pp. 1027 – 1045.
- [33]. Attia, H. A., (2007). On the effectiveness of variation in the physical variables on the generalized couette flow with heat transfer in a porous medium, *Research Journal of Physics*, Vol. I, No. 1, pp. 1 – 9.
- [34]. Makinde, O. D. and Osalusi, E., (2006). MHD steady flow in a channel filled with slip at permeable boundaries, *Romanian J. Physics*, Vol. 51, No. 3 – 4, pp. 319 – 328.
- [35]. Narahari, M., (2010). Effects of thermal radiation and free convection currents on the unsteady couette flow between two vertical parallel plates with constant heat flux at one boundary, *WSEAS Transactions on Heat and Mass Transfer*, Issue 1., Vol. 5, pp. 21 – 30.
- [36]. Israel – Cooke, C., Amos, E. and Nwaigwe, C., 2010. MHD oscillatory Couette flow of a radiating viscous fluid in a porous medium with periodic wall temperature, *Am. J. Sci. Ind. Res.*, Vol. 1, No. 2, pp. 326 – 331, doi:10.5251/ajsir.2010.1.2.326.331
- [37]. Brewster, M. Q., (1992). *Thermal radiative transfer & properties*, John Wiley & Sons.
- [38]. Antia, M., (1991). *Numerical Methods for Scientists and Engineers*, Tata McGraw – Hill, New Delhi.